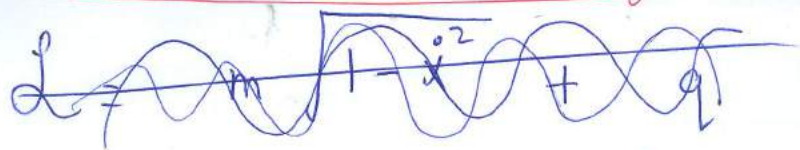


Relativistic Equation of motion

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Let's start with the action for a free particle.

$$S = -m \int_{t_a}^{t_b} d\tau = -m \int_{t_a}^{t_b} \sqrt{1 - \dot{x}^2} dt$$

where $\dot{x}^2 = |\dot{\vec{x}}|^2$.

For the interaction, let's construct a Lorentz invariant term

$$q \int_a^b dx^\mu A_\mu(t, \vec{x})$$

where a and b are two points (end-points) in Minkowski space.

$$\begin{aligned} S &= -m \int_{t_a}^{t_b} \sqrt{1 - \dot{x}^2} dt + q \int_a^b dx^\mu A_\mu(t, \vec{x}) \\ &= -m \int_{t_a}^{t_b} \sqrt{1 - \dot{x}^2} dt + q \int_{t_a}^{t_b} dt \frac{dx^\mu}{dt} A_\mu(t, \vec{x}) \end{aligned}$$

$$\begin{aligned} \Rightarrow \mathcal{L} &= -m \sqrt{1 - \dot{x}^2} + q v^\mu A_\mu(t, \vec{x}); \quad \mu = 0, 1, 2, 3. \\ &= -m \sqrt{1 - \dot{x}^2} + q v^0 A_0(t, \vec{x}) + q v^i A_i(t, \vec{x}) \end{aligned}$$

plus sign

Euler-Lagrange equation for the particle is :

Go

$$\frac{d}{dt} \left(\frac{\partial \mathcal{L}}{\partial \dot{x}^p} \right) = \frac{\partial \mathcal{L}}{\partial x^p} \quad \text{for each } p = 1, 2, 3.$$

Now let's look at these terms, one by one.

$$\begin{aligned} \frac{\partial \mathcal{L}}{\partial \dot{x}^p} &= \frac{\partial}{\partial \dot{x}^p} \left(-m \sqrt{1 - \dot{x}^p \dot{x}_p} \right) + q A_p(t, \vec{x}) \\ &= -\frac{m}{2\sqrt{1 - \dot{x}^2}} \left(-\frac{\partial}{\partial \dot{x}_p} (\dot{x}^p \dot{x}_p) \right) + q A_p(t, \vec{x}) \\ &= \frac{m}{2\sqrt{1 - \dot{x}^2}} \frac{\partial}{\partial \dot{x}^p} (\dot{x}^p (-\dot{x}^p)) + q A_p(t, \vec{x}) \\ &= -\frac{m}{\sqrt{1 - \dot{x}^2}} \dot{x}^p + q A_p(t, \vec{x}) \\ &= +\frac{m}{\sqrt{1 - \dot{x}^2}} \dot{x}_p + q A_p(t, \vec{x}) \end{aligned}$$

$$\text{LHS} = \frac{d}{dt} \left(\frac{m \dot{x}_p}{\sqrt{1 - \dot{x}^2}} \right) + q \frac{d}{dt} (A_p(t, \vec{x})) \quad \text{note the change from } p \text{ to } n.$$

$$\begin{aligned} \text{RHS} = \frac{\partial \mathcal{L}}{\partial x^p} &= q \underline{v}^0 \frac{\partial}{\partial x^p} A_0(t, \vec{x}) \\ &+ q \underline{v}^n \frac{\partial}{\partial x^p} A_n(t, \vec{x}) \end{aligned}$$

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$$RHS = q v^0 \frac{\partial}{\partial x^p} A_0 + q v^n \frac{\partial}{\partial x^p} A_n$$

(I am suppressing the argument (t, \vec{x}) on the R.H.S for the time being).

Hence

$$\begin{aligned} \frac{d}{dt} \left(\frac{m \dot{x}_p}{\sqrt{1 - \dot{x}^2}} \right) + q \frac{d}{dt} (A_p(t, \vec{x})) \\ = q v^0 \frac{\partial}{\partial x^p} A_0 + q v^n \frac{\partial}{\partial x^p} A_n \end{aligned}$$

Investigate $\frac{d}{dt} (A_p(t, \vec{x}))$ a bit closely.

$$\frac{d}{dt} (A_p(t, \vec{x})) = \frac{\partial A_p}{\partial t} + \dot{x}^n \frac{\partial A_p}{\partial x^n}$$

$$n = 1, 2, 3.$$

$$\begin{aligned} \therefore \frac{d}{dt} \left(\frac{m \dot{x}_p}{\sqrt{1 - \dot{x}^2}} \right) + q \frac{\partial A_p}{\partial t} + q \dot{x}^n \frac{\partial A_p}{\partial x^n} \\ = q v^0 \frac{\partial}{\partial x^p} A_0 + q v^n \frac{\partial}{\partial x^p} A_n \end{aligned}$$

$$\left[\frac{dx^0}{dt} = \frac{dt}{dt} = 1 \right]$$

Ob

$$\frac{d}{dt} \left(\frac{m \dot{x}_p}{\sqrt{1-\dot{x}^2}} \right) = q \dot{x}^0 \frac{\partial A_0}{\partial x^p} - q \frac{\partial A_p}{\partial t} + q \dot{x}^n \left(\frac{\partial A_n}{\partial x^p} - \frac{\partial A_p}{\partial x^n} \right)$$

$$\frac{d}{dt} \left(\frac{m \dot{x}_p}{\sqrt{1-\dot{x}^2}} \right) = q \left(\frac{\partial A_0}{\partial x^p} - \frac{\partial A_p}{\partial t} \right) + q \dot{x}^n \left(\frac{\partial A_n}{\partial x^p} - \frac{\partial A_p}{\partial x^n} \right)$$

Eq. ①

Define $\frac{\partial A_0}{\partial x^p} - \frac{\partial A_p}{\partial t} = \frac{\partial A_0}{\partial x^p} - \frac{\partial A_p}{\partial x^0} = E_p$

e.g. $E_x = \frac{\partial A_0}{\partial x} - \frac{\partial A_x}{\partial t}$ etc. where $A_0 = -\phi$

and ϕ is the conventional electric potential.

Let's identify the second term on the RHS.

Now

$$\left(\vec{v} \times \vec{B} \right)_p = \epsilon_{pqn} \dot{x}^q B_n$$

If $\vec{B} = \vec{\nabla} \times \vec{A}$ and $B_n = \epsilon_{nst} \frac{\partial}{\partial x^s} A_t$

(Def. of curl in tensor notation).

then $\left(\vec{v} \times \vec{B} \right)_p = \epsilon_{pqn} \dot{x}^q \epsilon_{nst} \frac{\partial A_t}{\partial x^s}$

$$(\vec{v} \times \vec{B})_p = \epsilon_{pqz} \epsilon_{rst} v^z \frac{\partial A_t}{\partial x^r}$$

$$= \epsilon_{npq} \epsilon_{rst} v^z \frac{\partial A_t}{\partial x^r}$$

$$= (\delta_{ps} \delta_{qt} - \delta_{pt} \delta_{qs}) v^z \frac{\partial A_t}{\partial x^r}$$

(using an identity involving Levi-Civita symbols)

$$= v^t \frac{\partial A_t}{\partial x^p} - v^z \frac{\partial A_p}{\partial x^z}$$

refer to your electrodynamics course)

$$= v^t \frac{\partial A_t}{\partial x^p} - v^t \frac{\partial A_p}{\partial x^t}$$

$$= v^n \frac{\partial A_n}{\partial x^p} - v^n \frac{\partial A_p}{\partial x^n}$$

(change variables in the repeated sum)

$$= \dot{x}^n \left(\frac{\partial A_n}{\partial x^p} - \frac{\partial A_p}{\partial x^n} \right)$$

which is identical to the second term on the RHS.

Hence

$$\boxed{\frac{d}{dt} \left(\frac{m \dot{x}_p}{\sqrt{1 - \dot{x}^2}} \right) = q E_p + q (\vec{v} \times \vec{B})_p}$$

which is the correct relativistic equation for the 3-vector. Let's cast this in a manifestly Lorentz invariant form.

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Multiply by $\frac{dt}{d\tau} = \frac{dx^0}{d\tau}$

$n = 1, 2, 3$
 $p = 1, 2, 3$

$$\begin{aligned}
 &= q \left(\frac{dx^0}{d\tau} \frac{\partial A_0}{\partial x^p} + \frac{dx^n}{d\tau} \frac{\partial A_n}{\partial x^p} \right) \\
 &\quad - q \left(\frac{dx^0}{d\tau} \frac{\partial A_p}{\partial x^0} + \frac{dx^n}{d\tau} \frac{\partial A_p}{\partial x^n} \right) \\
 &= q \left(\frac{dx^\mu}{d\tau} \frac{\partial A_\mu}{\partial x^p} \right) - q \left(\frac{dx^\mu}{d\tau} \frac{\partial A_p}{\partial x^\mu} \right)
 \end{aligned}$$

Hence

p is a free index and $p = 1, 2, 3$. μ is a contracted index.

Let's replace μ by v , no harm done.

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$$\frac{d}{d\tau} (m u_p) = q \frac{dx^v}{d\tau} \left(\frac{\partial A_v}{\partial x^p} - \frac{\partial A_p}{\partial x^v} \right) \quad \text{--- (2)}$$

Now S is Lorentz invariant. Equations of motion are Lorentz invariant.
 If the three spatial equations of motion derived from L (a Lorentz invariant Lagrangian)

give three components of a 3-vector, each one being equal to the three components of another 3-vector (Eq. 2), then the time component must also be equal.

Eq.

Hence if I set the zeroth component in the left and R.H.S. of Eq. 2, then it should be a valid equality.

i.e.

$$\frac{d}{d\tau} (m u_0) = q \frac{dx^v}{d\tau} \left(\frac{\partial A_v}{\partial x^0} - \frac{\partial A_0}{\partial x^v} \right)$$

check.

$$\text{LHS} = \frac{d}{dt} \left(m \frac{1}{\sqrt{1-v^2}} \right) = m \frac{d}{dt} (u_0) = \frac{d}{dt} (\text{energy})$$

since $m u_0 = \text{energy}$ (zeroth term = the 4-velocity)

$$\text{RHS} = q \frac{dx^v}{d\tau} \left(\frac{\partial A_v}{\partial x^0} - \frac{\partial A_0}{\partial x^v} \right) = \cancel{q \frac{dx^v}{d\tau} E_v} =$$

$$\begin{aligned}
 &= q \left[\frac{dx^0}{d\tau} \left(\overset{\nearrow}{\frac{\partial A_0}{\partial x^0}} - \overset{\text{zero}}{\frac{\partial A_0}{\partial x^0}} \right) \right. \\
 &\quad \left. + \frac{dx^n}{d\tau} \left(\frac{\partial A_n}{\partial x^0} - \frac{\partial A_0}{\partial x^n} \right) \right] \\
 &= q \frac{dx^n}{d\tau} \left(-E_n \right)
 \end{aligned}$$

$$\frac{d}{d\tau} (mu_0) = -q \frac{dx^n}{d\tau} E_n$$

$$\frac{d\tau}{dt} \frac{d}{d\tau} (mu_0) = -q \frac{d\tau}{dt} \frac{dx^n}{d\tau} E_n$$

$$\frac{d}{dt} (mu_0) = -q \frac{dx^n}{dt} E_n$$

$$\text{rate of change of K.E.} = -q \vec{v} \cdot \vec{E} \quad \left(\begin{array}{l} \text{relativistic} \\ \text{work} \\ \text{energy principle} \end{array} \right)$$

Key idea

$\mathcal{L} \rightarrow$ Lorentz invariant

\Rightarrow equations of motion are Lorentz invariant

\Rightarrow space ³-vector component equations are formed

\Rightarrow the time component of the equation of motion will follow.

Because of the Lorentz invariance of the \mathcal{L}
equation of motion, Eq. (2) will also hold if I
replaced p_i with another Greek free index μ
which can take up the value 0, 1, 2, 3. We have
already verified it for $\mu = 1, 2, 3$ (Eq. (2))
and $\mu = 0$ follows because of Lorentz invariance of the
overall action. Besides, it makes sense intuitively as
it reproduces the relativistic work-energy principle.